

# The composition of Barium stars and the s process in AGB stars

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# Introduction – Ba stars

- G–K giants + dwarfs,  $[\text{Fe}/\text{H}] > -1.0$
- strong spectral features: carbon molecular bands + s-process elements
- radial velocity variation → binary systems
  - not intrinsic overabundance!
  - mass transfer
  - test: AGB s-process nucleosynthesis

# Introduction – Ba stars

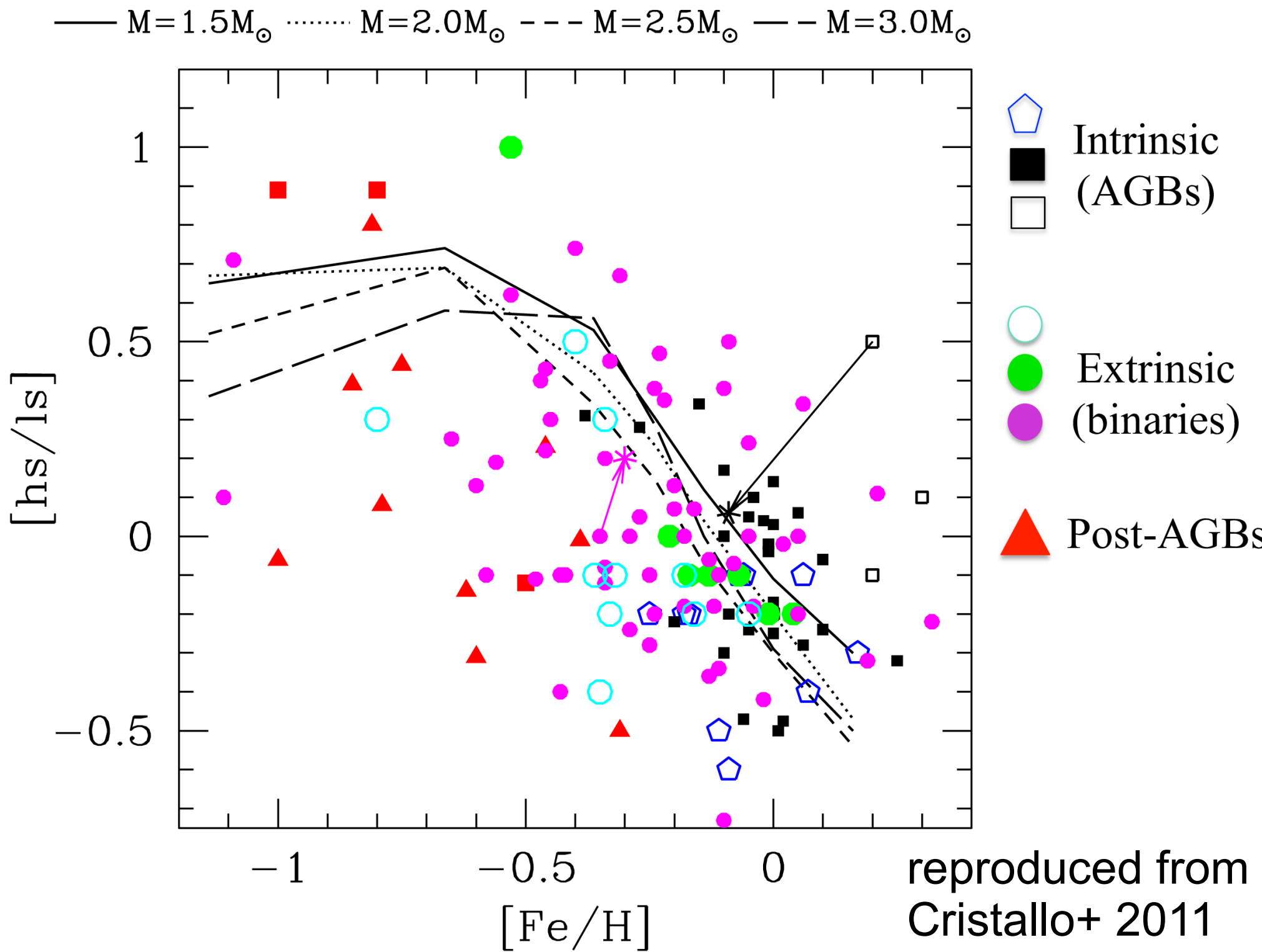
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- $[hs/l_s] \leftarrow s=? \quad hs/l_s=?$   
(2./1. peak)

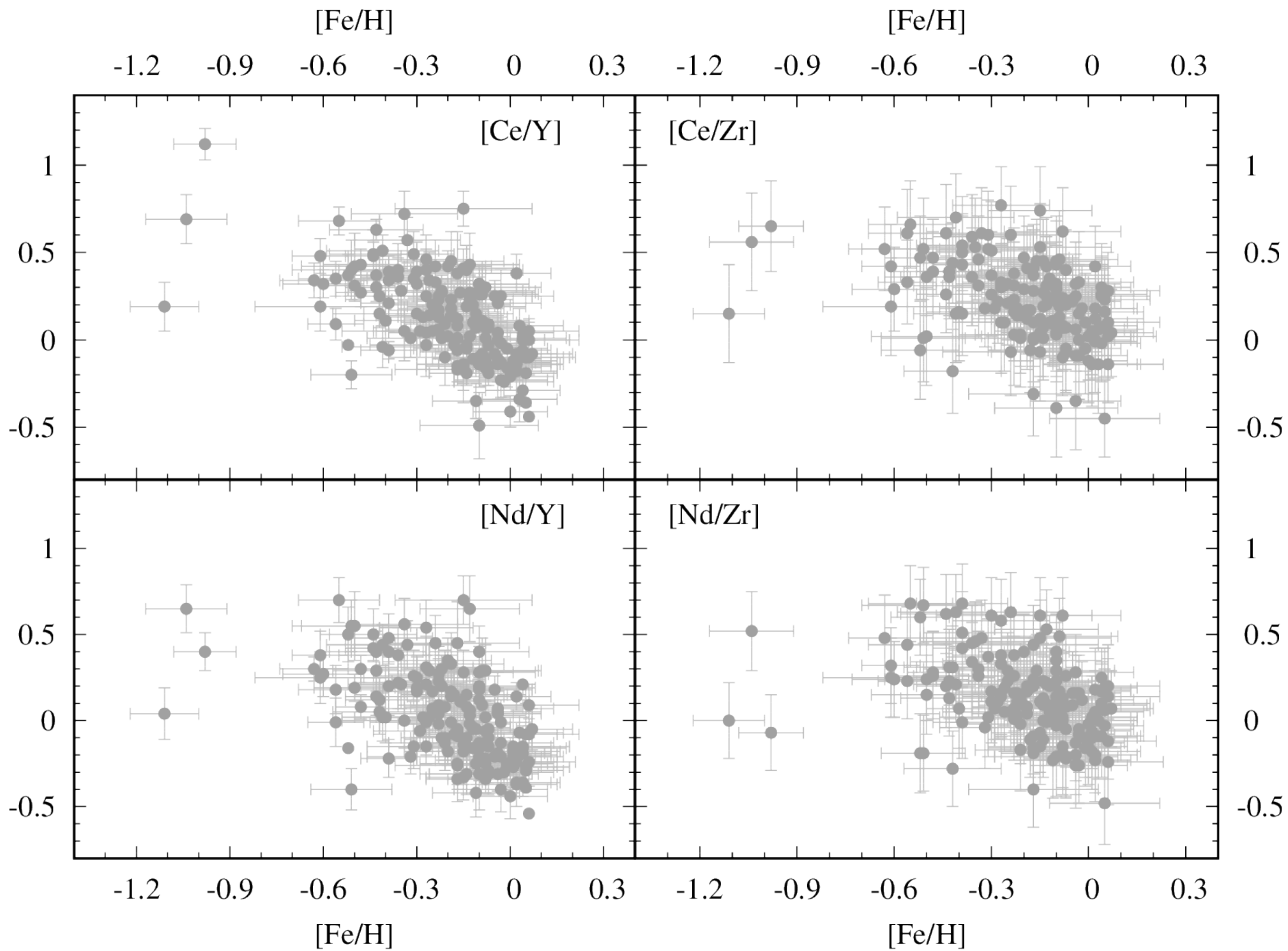
# Sample data

- e. g.: Yang+ (2016): 19 stars, Allen&Barbuy (2006): 26 stars, Antipova+ (2004): 16 stars
- de Castro et al. (2016) sample:
  - 182 giants (certain, candidate)
  - high resolution spectra (FEROS,  $R = 48000$ )
  - wide range in  $T_{\text{eff}}$  (4100-5400 K), mass (1-6  $M_{\text{Sun}}$ ), metallicity
  - Ba star? here: if  $[s/\text{Fe}] \geq 0.25 \rightarrow 13$  stars rejected ( $s = \text{Y, Zr, La, Ce, Nd}$ )

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- estimated error  $\rightarrow$  first time proper error calculations  
 **$[\text{hs/ls}] \rightarrow [\text{Ce/Y}], \dots$**





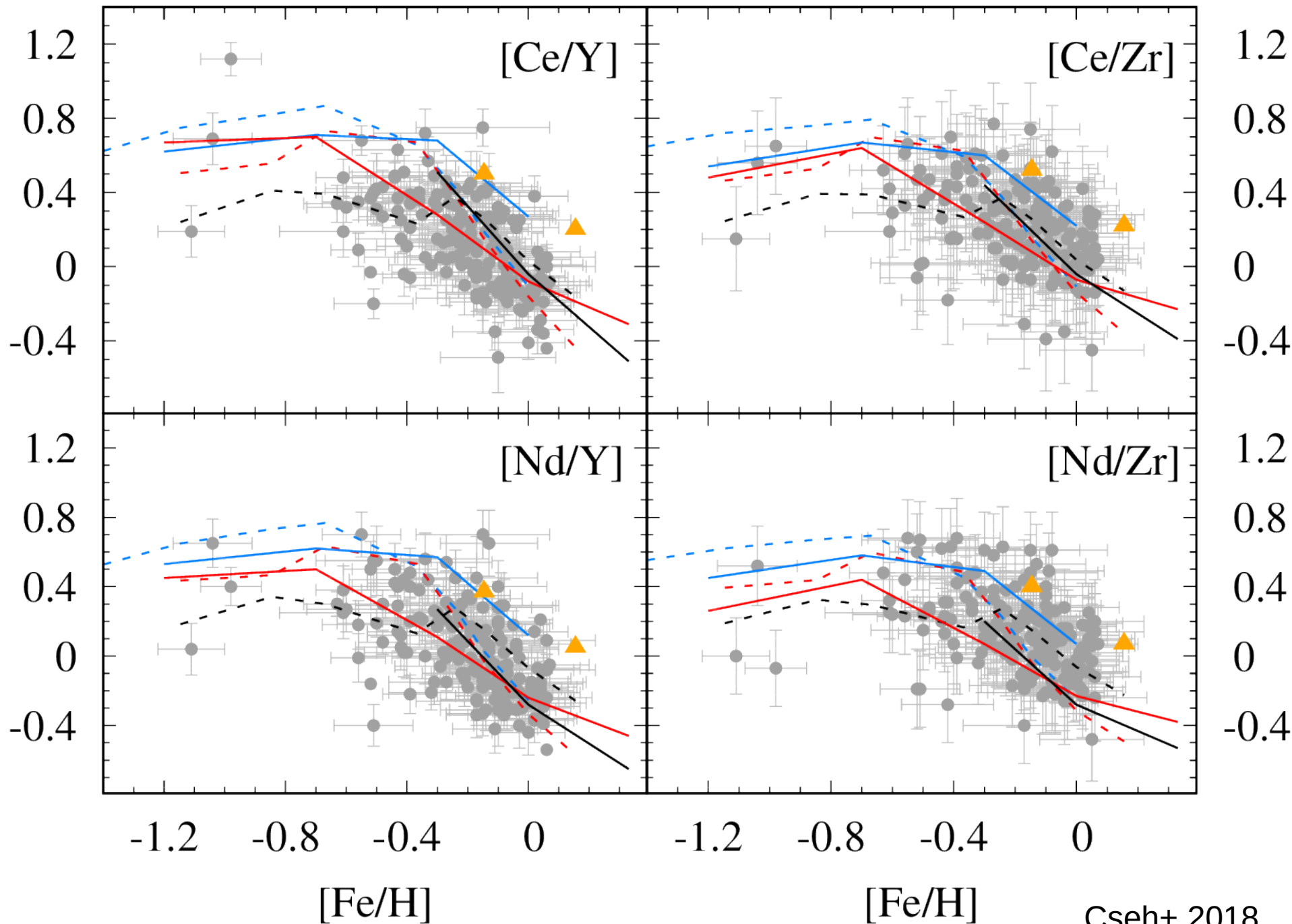
# Model comparison

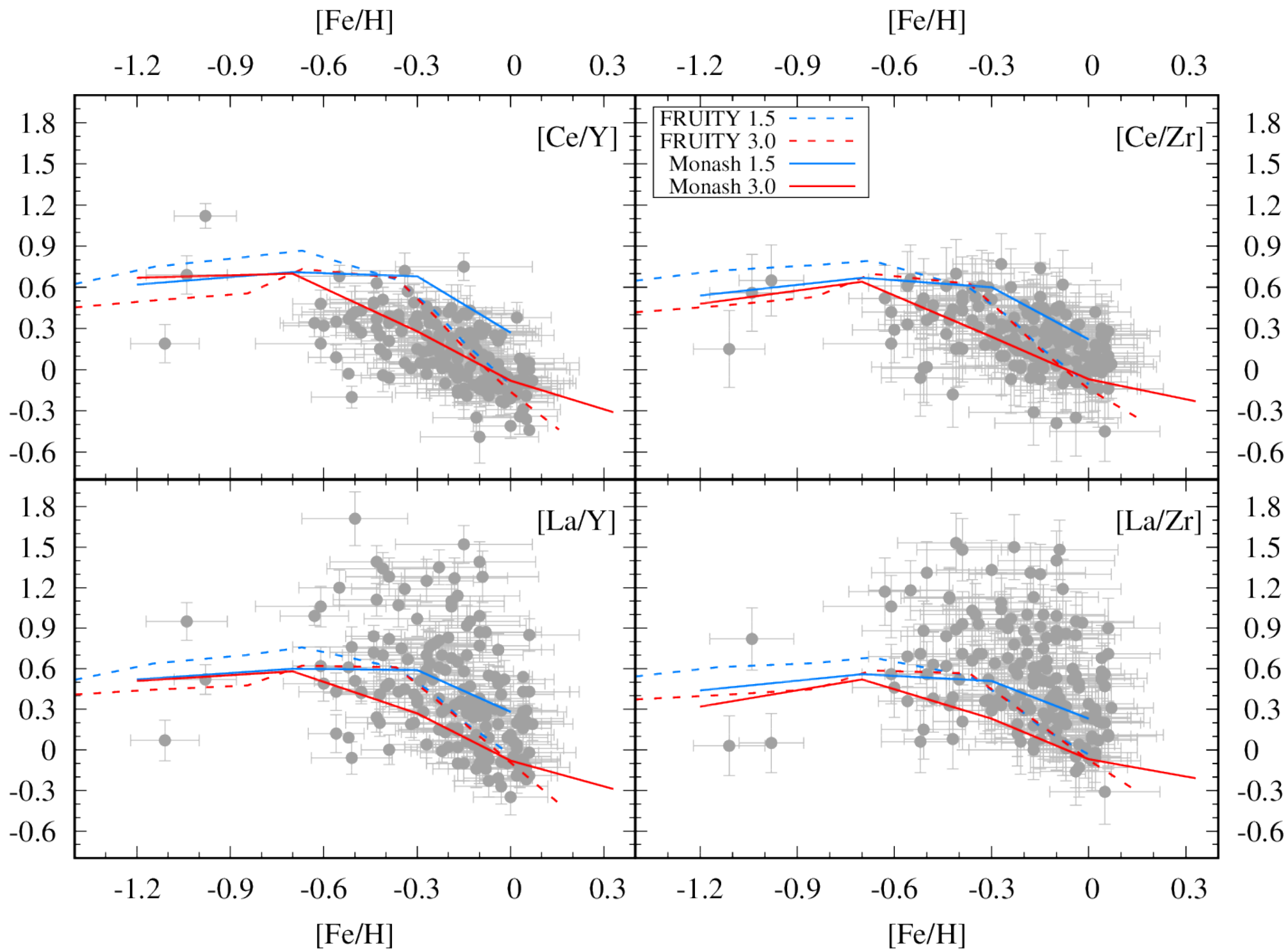
- final surface composition, without rotation,  $[s/Fe] \geq 0.25$   
(s = Y, Zr, La, Ce, Nd)
- different metallicities, masses
- **FRUITY + Monash + NuGrid: Battino+ 2016**

Cristallo+ 2016,  
Cristallo+ 2015,  
Straniero+ 2014,  
Piersanti+ 2013,  
Cristallo+ 2011,  
Cristallo+ 2009

Karakas+ 2018,  
Karakas & Lugaro 2016,  
Fishlock+ 2013

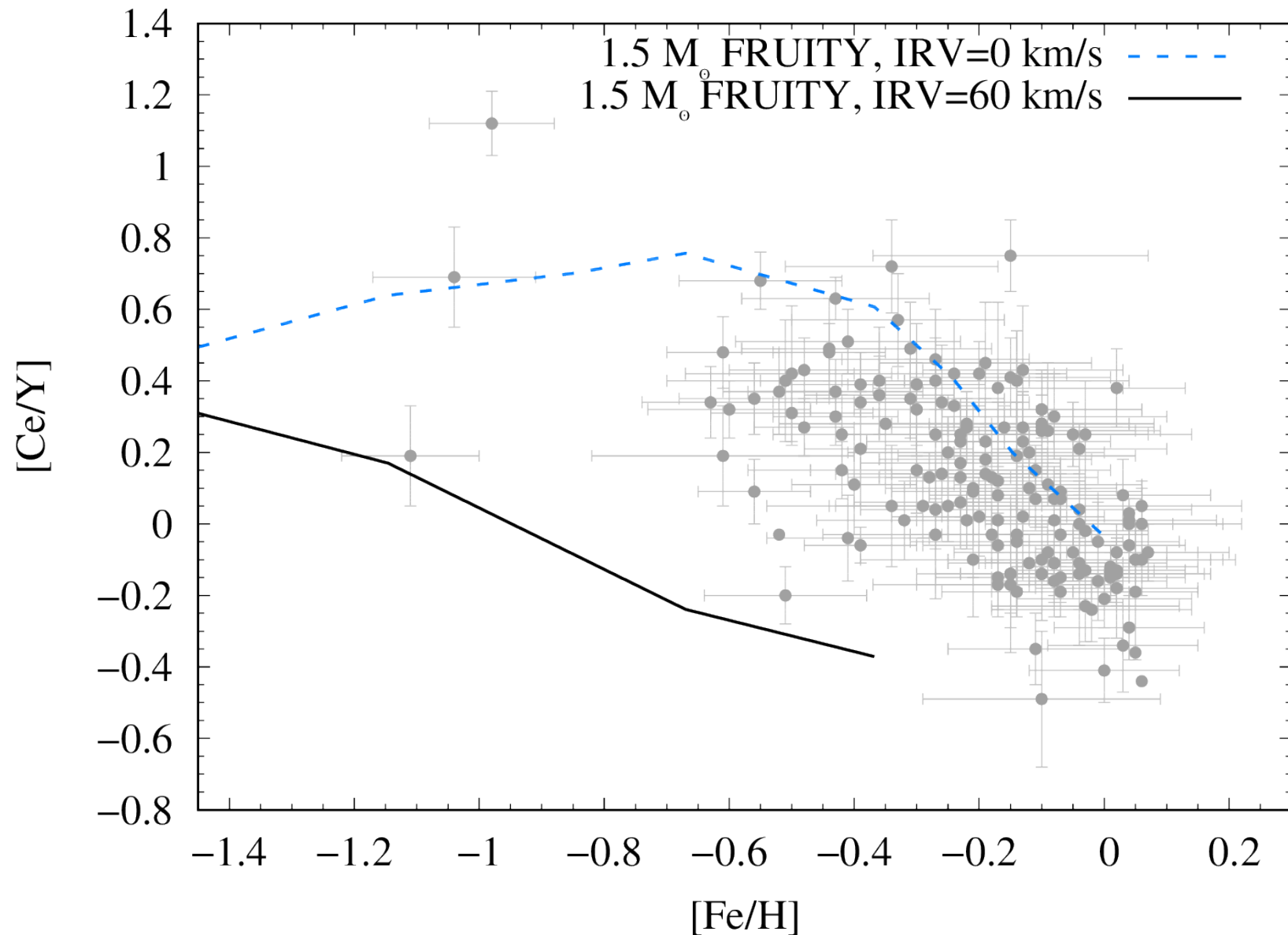
FRUITY 1.5M<sub>⊙</sub>    - - - -    FRUITY 4.0M<sub>⊙</sub>    - - - -    Monash 1.5M<sub>⊙</sub>    ————    Monash 4.0M<sub>⊙</sub>    ————  
 FRUITY 3.0M<sub>⊙</sub>    - - - -    NuGrid 3.0M<sub>⊙</sub>    ▲    Monash 3.0M<sub>⊙</sub>    ————    Monash 3.0M<sub>⊙</sub>    ————





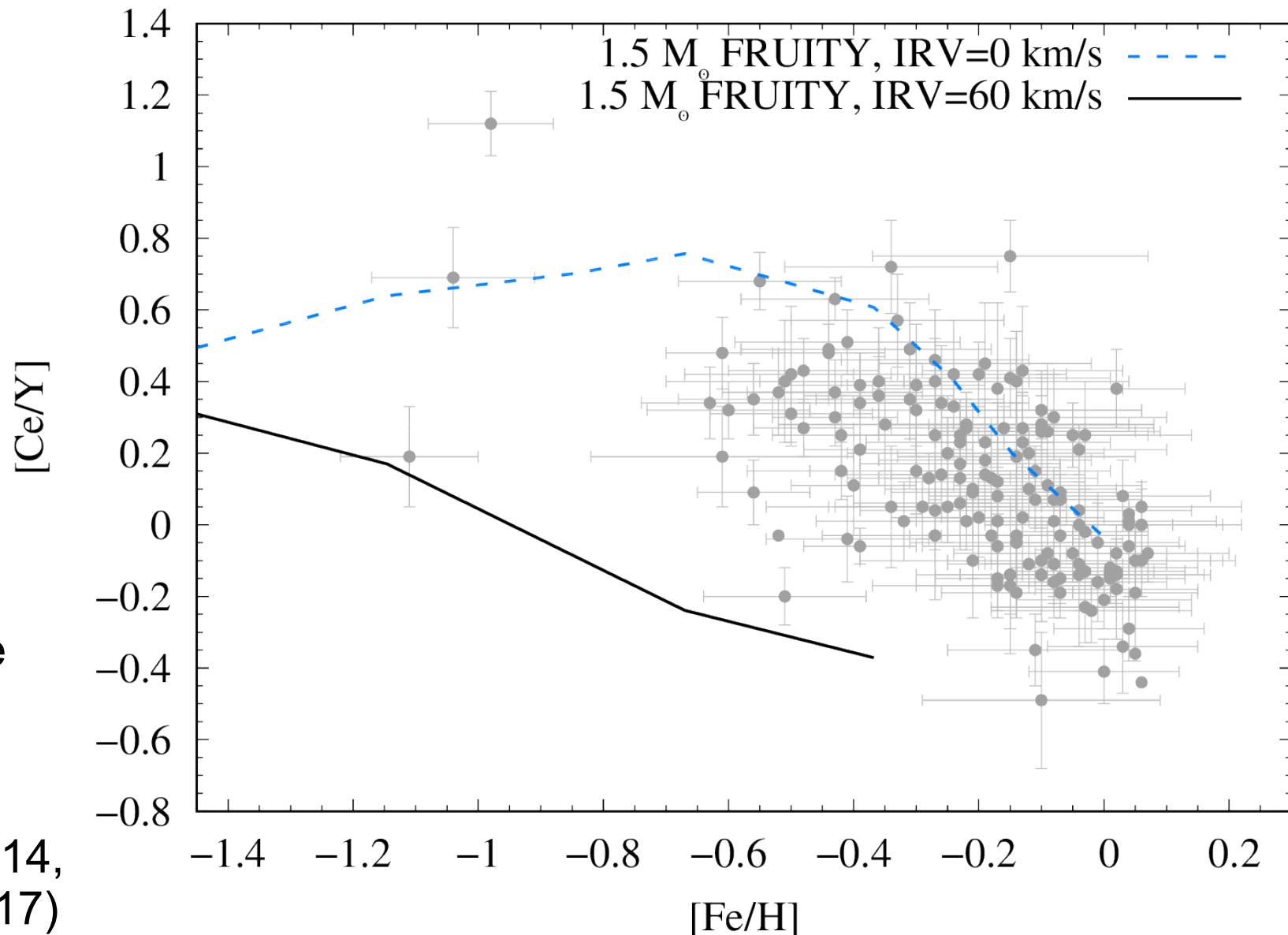
# ...with rotation

Cseh+ 2018



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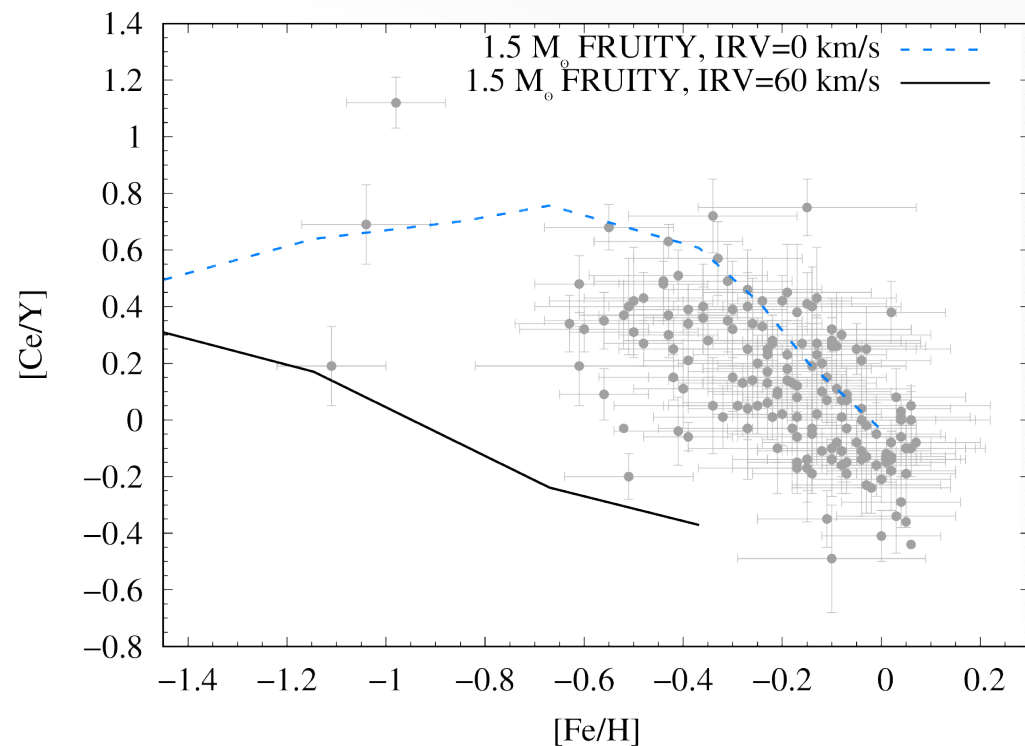
Cseh+ 2018



- rotation: minor effect
  - coupling: to slow down the core in giants
  - agreement with asteros.
- (Cantiello+ 2014, Hermes+ 2017)

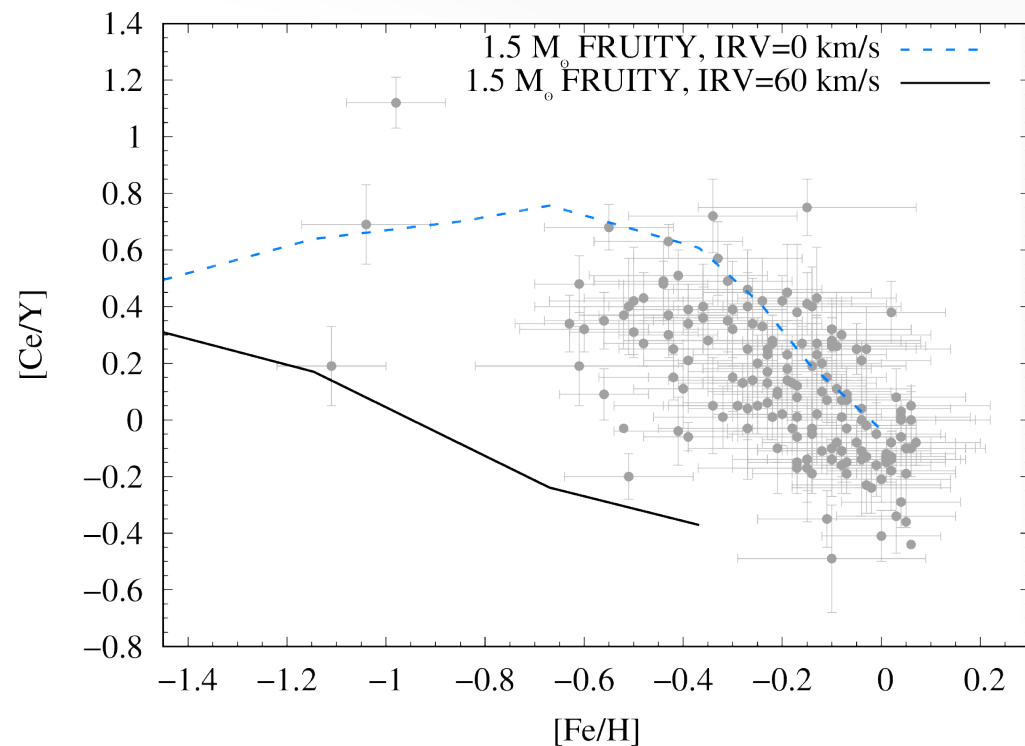
# Summary

- new Ba star observations  
→ split elements, own error bar
- trend agrees with models →  
main n source:  $^{13}\text{C}$  nuclei
- rotational mixing: minor effect  
→ in agreement with asteroseismology  
(Cantiello+ 2014, Hermes+ 2017)



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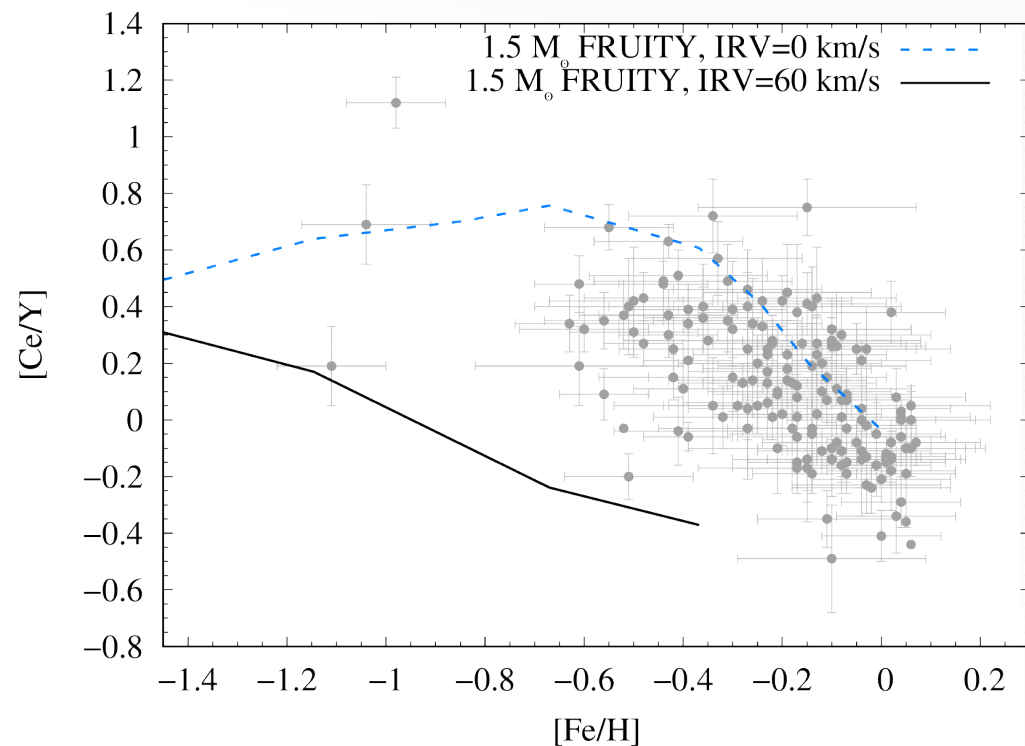


future plans (work in progress):

- other elements, individual stars
- higher metallicity sample
- comparison with other s-process enhanced stars

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**Thank you for your attention!**